



Enabling Grids for E-scienceE

Energie noire et grille de calcul Le projet ZEN

A. Tilquin

Centre de Physique des Particules de Marseille

www.eu-egee.org



- Scientific Goal:

Address open questions of fundamental cosmology:
dark matter/dark energy sector and primordial universe
Construct a public tool....

- Main people involve:

ANR program

CPPM

Alain Bonnissent, Anne Ealet, Dominique Fouchez, Lei Sun, Diane Talon-Esmieu, Charling.
Tao

CEA

Philippe.Brax, Jean Batiste Melin, Christophe.Yèche, Dominique.Yvon, Nathalie Palanque-
Delabrouille, Alexandre.Réfrégier, Jim.Rich

CPT

Christian. Marinoni, Pierre Taxil, Jean-Marc Virey, Stefan. Linden

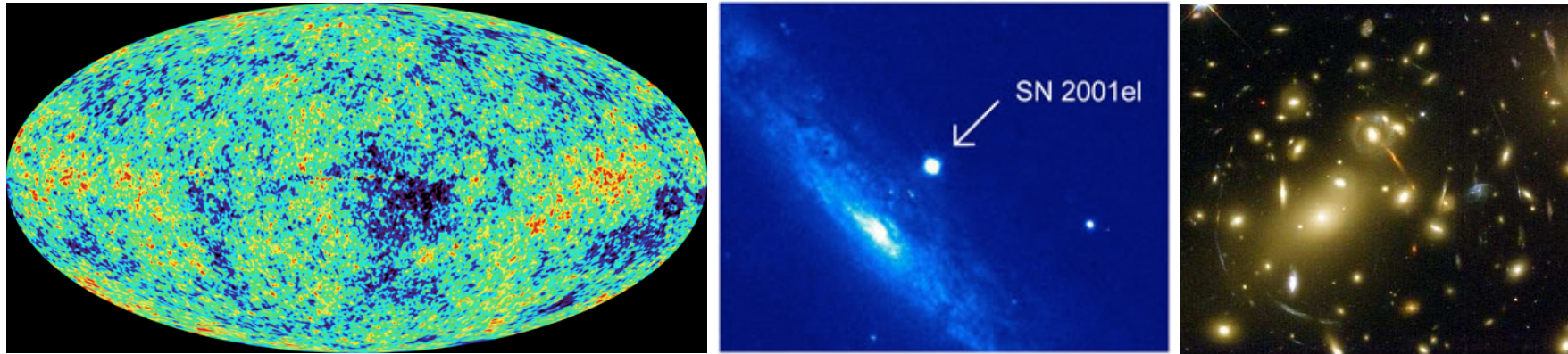
LAL

Francois Couchot, Olivier Perdereau, L. Perotto, S. Plaszczynski , C. Rosset

FCPPL: Franco-China-Particle-Physics-Laboratory program

IHEP/PKU/NAOC/Tsinghua U/Beijing Normal U.

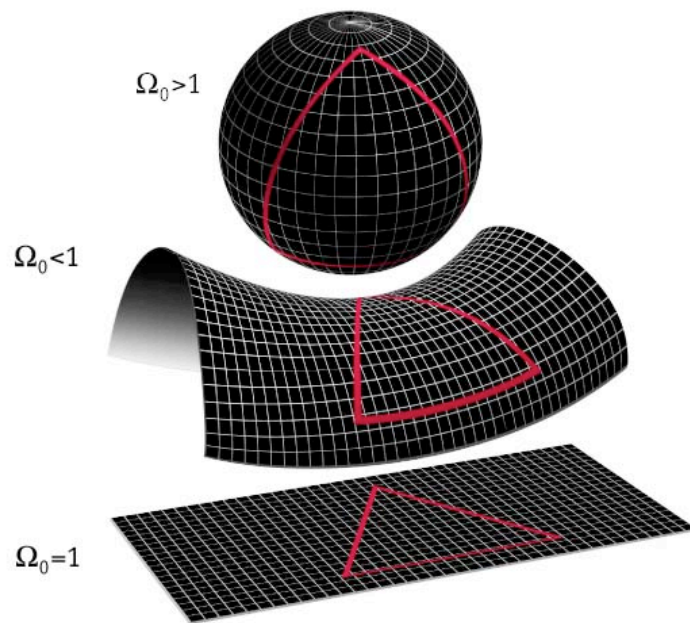
ZHANG XinMin, ZHAO Gongbo , QIAN Zuxuan , XIA Junqing, FAN Zuhui, LI Hong , Zhu
Zhonghong, QIN Bo, DENG Jinsong , CHAO Wu , ZHOU Xu , WANG Xiaofeng



- Few words about cosmology and dark energy
- How to characterize dark energy
- Statistical method and datagrid
- Results already obtained by the Chinese and French groups (ESR and Euchina)
- Prospectives

- Cosmology based on:
Homogenous and isotropic Universe

General relativity: $G_{\mu\nu} = 8\pi G T_{\mu\nu}$



MAP990006

Energy content: $\Omega = \rho/\rho_c$

Equation of state: $w(z)=p/\rho$

Matter (Ω_m) : $w = 0$

Radiation (Ω_r) : $w = 1/3$

Cosmological cste (Ω_Λ): $w=-1$

Dark energy (Ω_X) : $w(z)$

$$\Omega_T = \Omega_m + \Omega_r + \Omega_X$$

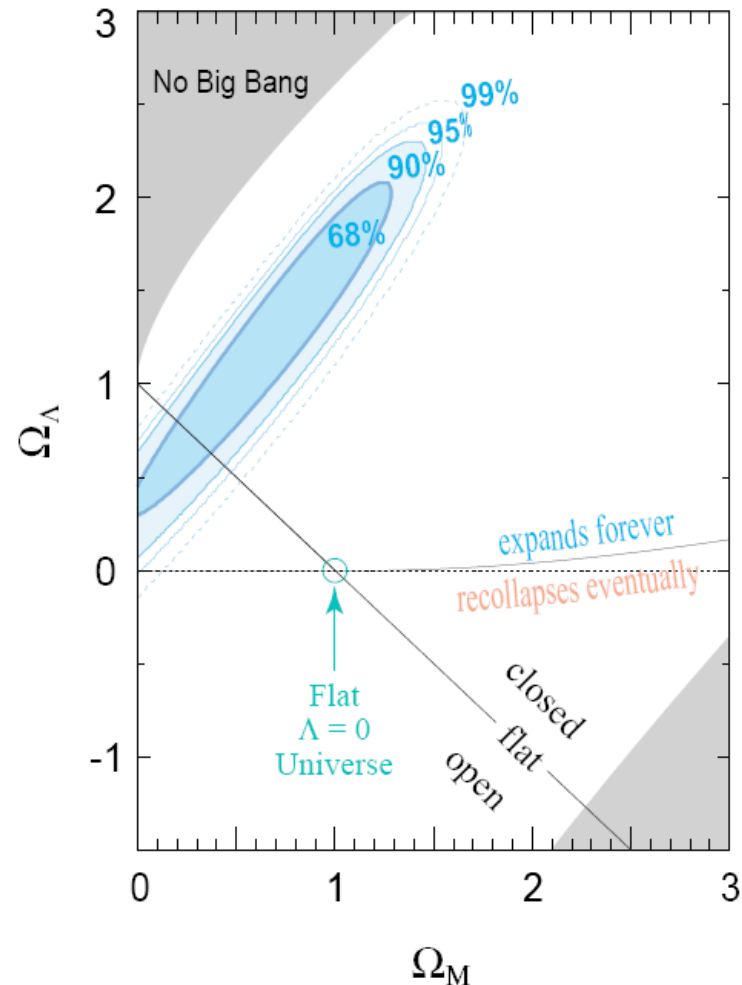
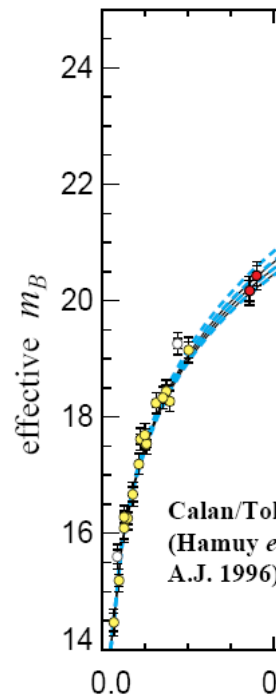
$(\rho_c=10^{-29} \text{ g/cm}^3)$

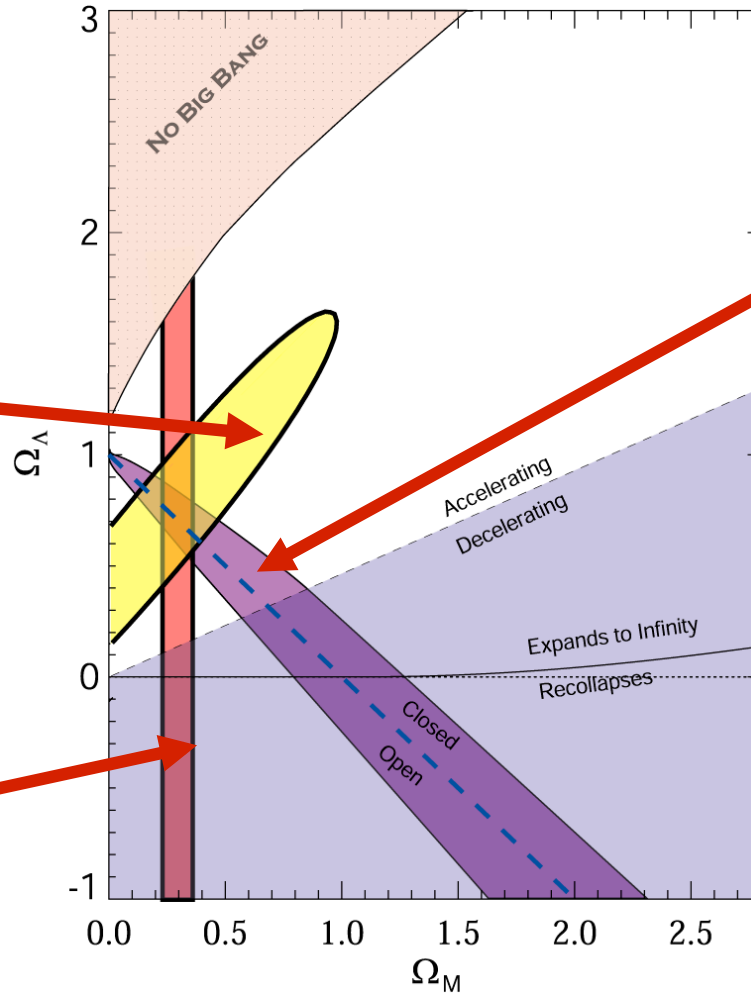
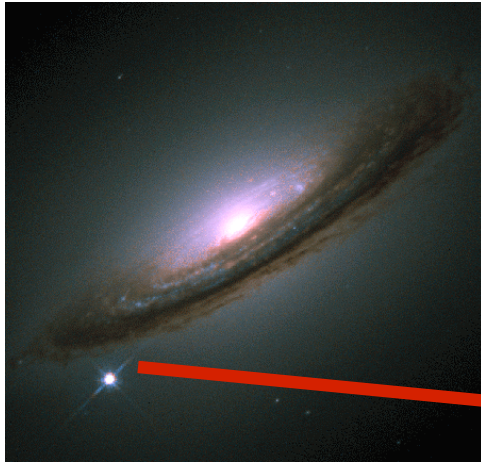
- In 1998, the Supernova Cosmology Project and High-z team shown that high red-shift supernovae are fainter than expected: a new energy component is needed.

Dark energy or cosmological constant characterized by reduce density: $\Omega_\Lambda = \rho_\Lambda / \rho_c$

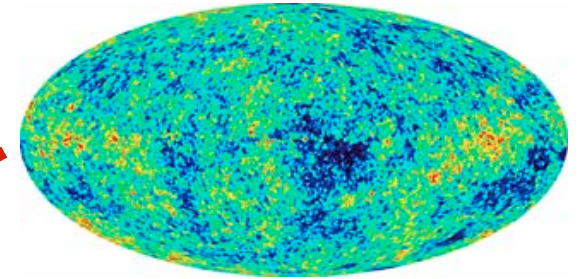
For a flat Universe:

$$\Omega_\Lambda = 0.72^{+0.09+0.05}_{-0.08-0.04}$$





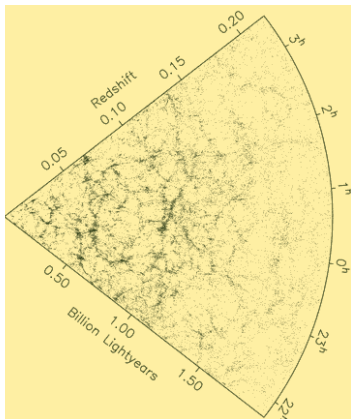
CMB(WMAP)



Universe is mainly dark

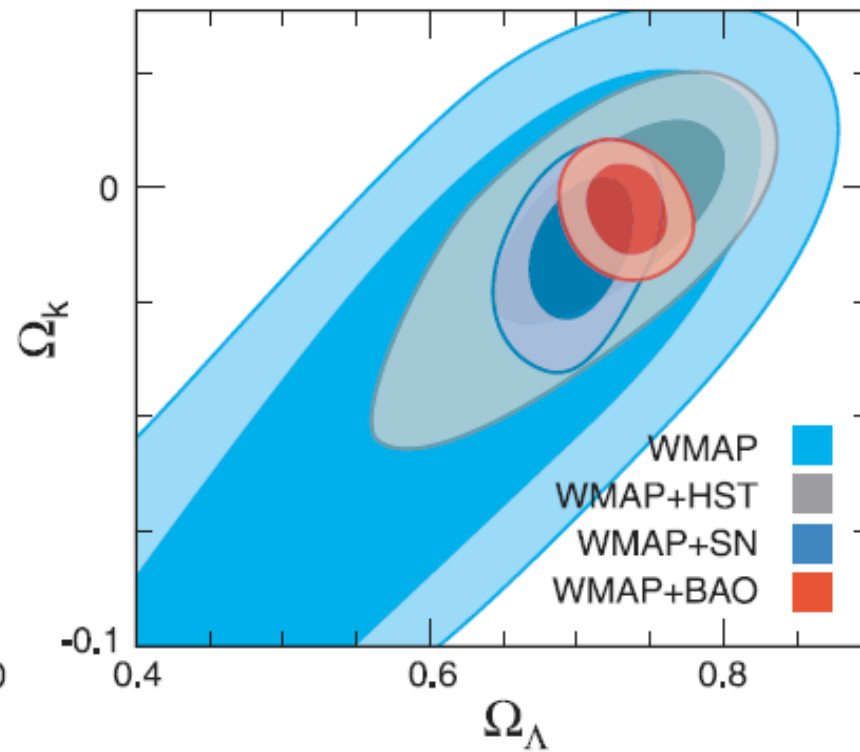
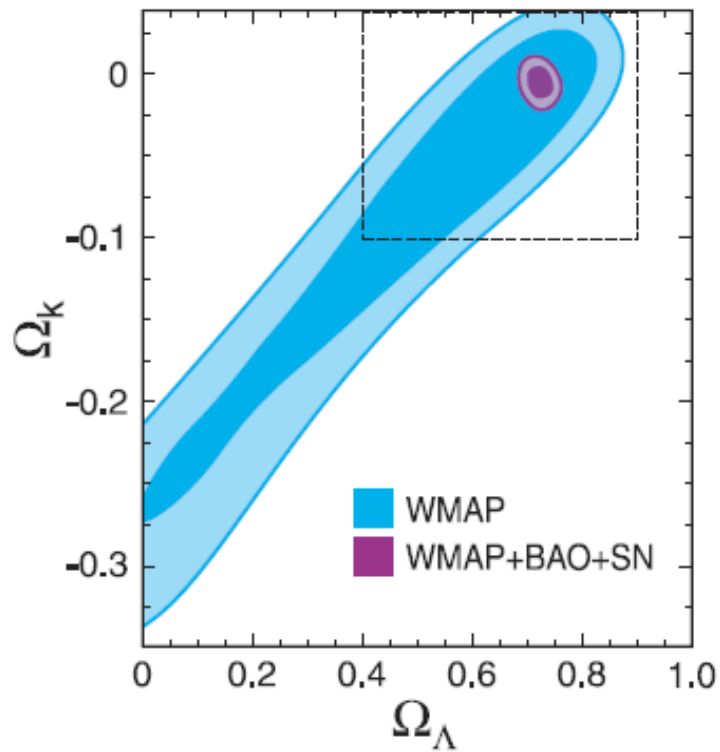
**~70% DE
and
~25% DM.**

LSS

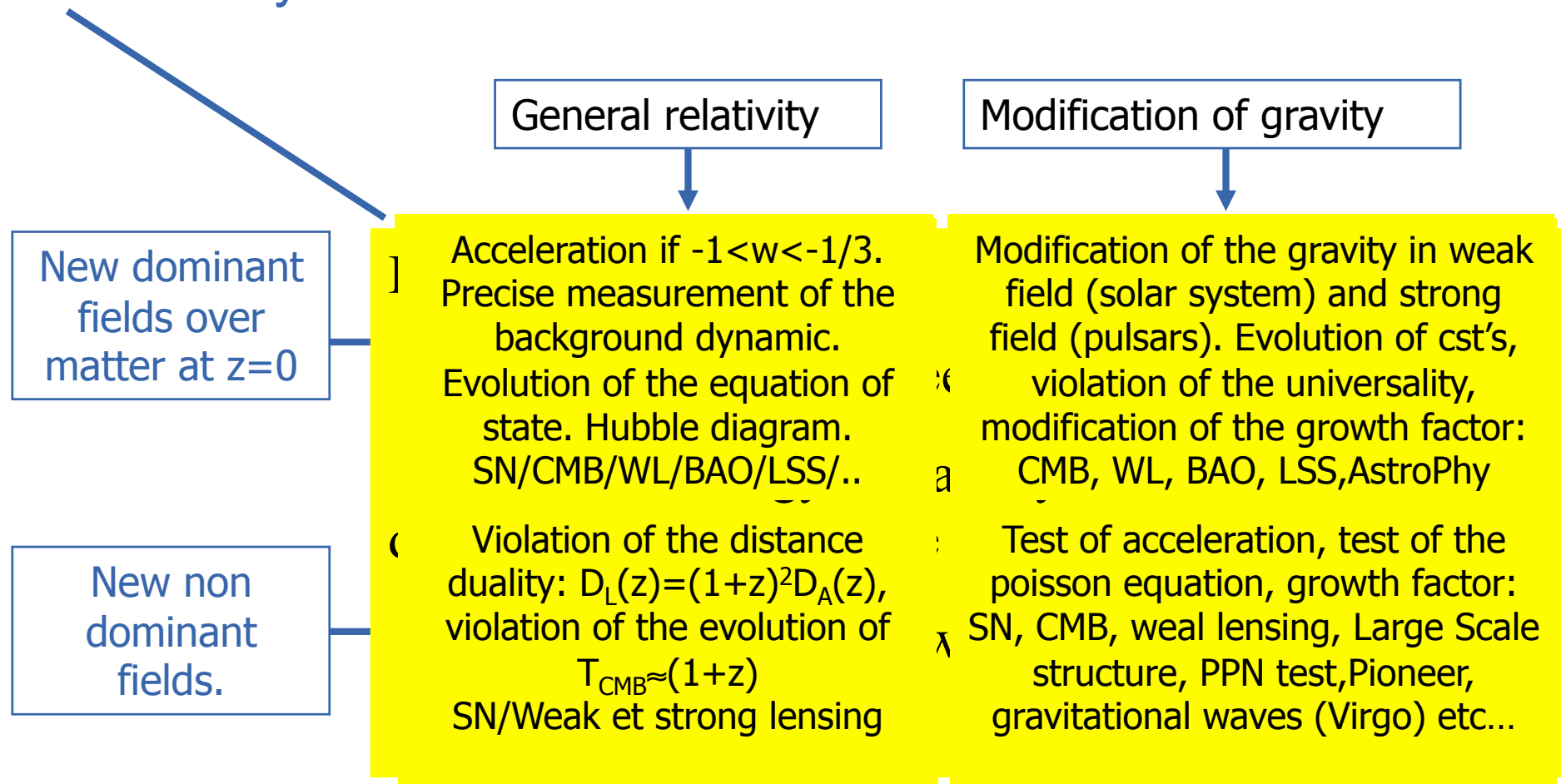


Our Univers is accelerating

a 2006 status



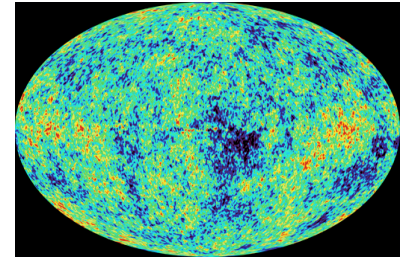
- From where this acceleration is coming from ?
- Two main classes of theory/model : Particle physics or Gravity ?



Degeneracies between parameters imply multi-probes analysis sensitive to complementary quantities:

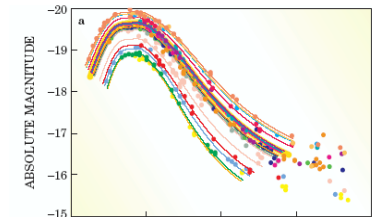
CMB

Snapshot at $\sim 400,000$ yr, viewed from $z=0$
 Angular diameter distance to $z \sim 1000$
 Growth rate of structure (from ISW)



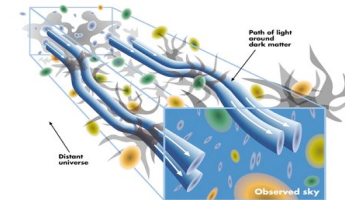
Supernovae

Standard candle
 Luminosity distance



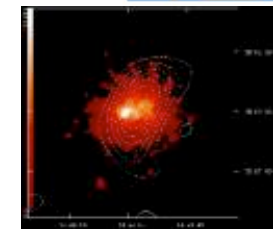
Cosmic Shear

Evolution of dark matter perturbations
 Angular diameter distance
 Growth rate of structure



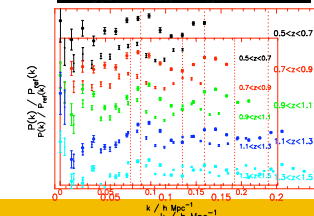
Cluster counts

Evolution of dark matter perturbations
 Angular diameter distance
 Growth rate of structure



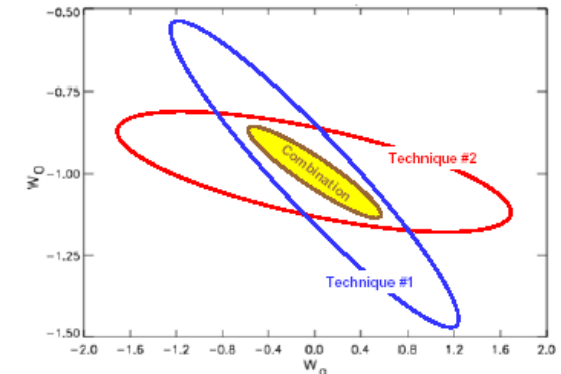
Baryon Wiggles

Angular diameter distance



- **Important number of parameters, cosmological and astrophysical**

- Ω_b/Ω_m density for baryon/matter
- Ω_ν density for neutrino's
- Ω_T curvature density
- H Hubble constant,
- n_s spectral index,
- τ reionisation optical depth
- σ_8 normalization for CMB, WL and BAO.
- m_s normalization for SNIa.
- y_{he} Helium fraction
- w_0, w_a Equation of state....



- **Efficient statistical tools needed:**

- Bayesian statistic and MCMC : Chinese and astrophysicist choice
- Frequentist statistic and datagrid : French and particle physicist choice

- **Statistic based on $\chi^2(\Omega_i, \mathbf{w}_0, \mathbf{w}_a, \dots)$**
- **Minimum using the gradient method:**

$$\frac{\partial \chi^2}{\partial \Omega_i} = 0$$

- **Numerical resolution and iterative:**

- **Error computation:** $U_{kl}^{-1} = \frac{1}{2} \left[\frac{\partial^2 \chi^2}{\partial \Omega_k \partial \Omega_l} \right]$

$$(\Omega_i - \Omega_i^o) = - \left(\frac{\partial^2 \chi^2}{\partial \Omega_k \partial \Omega_i} \Big|_{\Omega_k^o} \right)^{-1} \frac{\partial \chi^2}{\partial \Omega_k} \Big|_{\Omega_k^o}$$

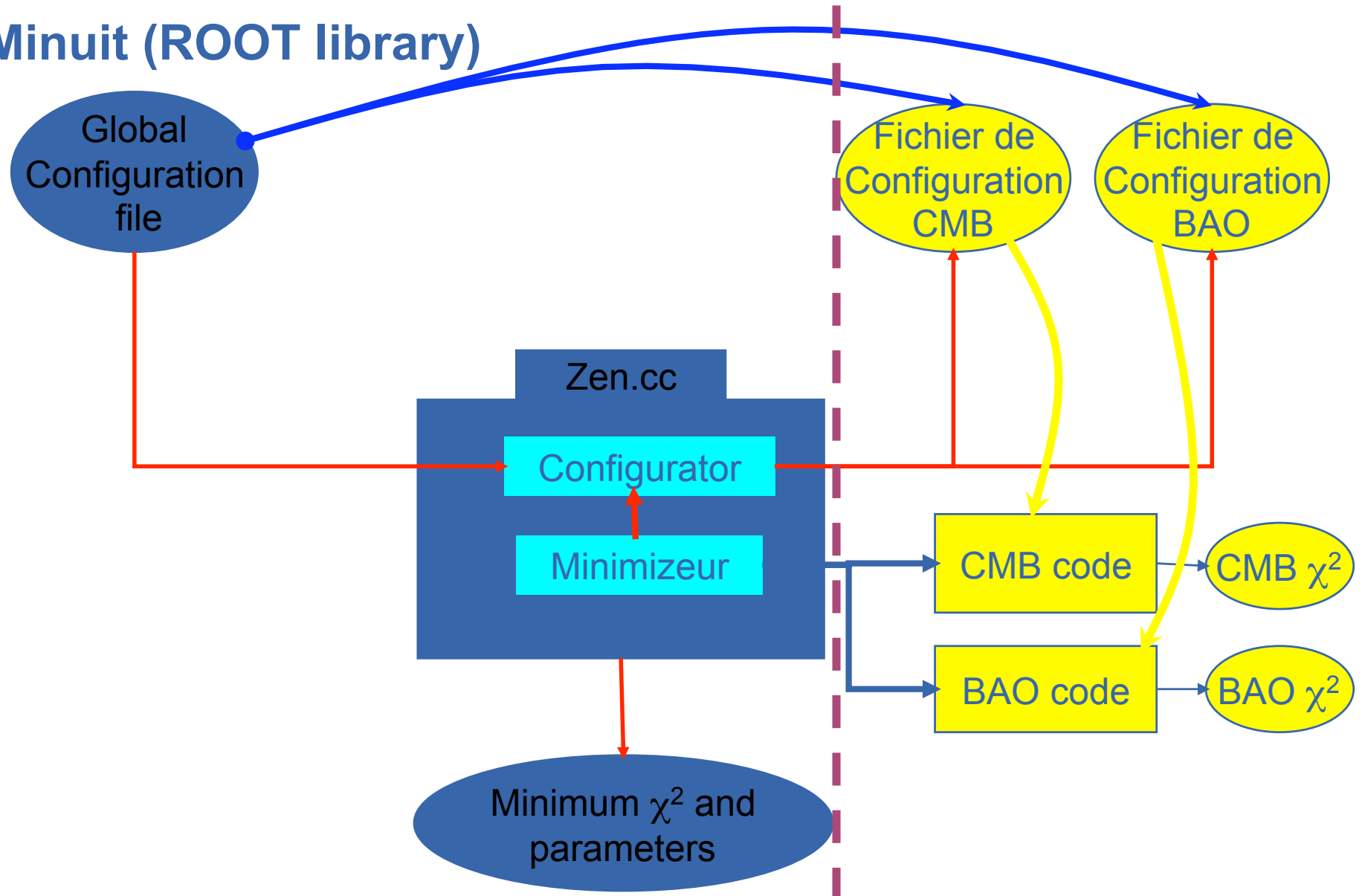
- **Contour: Solving the equation $\chi^2 = \chi^2_{\min} + \mathbf{s}^2$**
- **Marginalization obtained by minimization:**

$$\chi^2(\mathbf{w}_0, \mathbf{w}_a) = \chi^2_{\min}(\Omega_i | \mathbf{w}_0, \mathbf{w}_a)$$

The contour is constructed by minimizing the χ^2 on a grid of points (minimum 20*20) and iso- χ^2 are constructed using interpolation. Each point (20 hours of computing) is calculated on a CE.

**A simple contour requires about 1 year of CPU on a single CPU.
Thanks to datagrid (result in few days)**

Minuit (ROOT library)



```

action = compute_chi2/find_minimum
simulation = no
#-----
scan  = wa
from  = 0.
upto  = 0.1
nbscan = 100
# contour ?
scan2d = none
from2d = -1.0
upto2d = 0.0
nbscan2d = 2
#-----
use_sn = yes
use_bao = yes
use_wl = no
use_cmb = yes
use_hst = no
cmb_code = cmbeasy
flatness = yes
#-----

```

```

#-----
fit_omega_b = yes
fit_omega_cdm = yes
fit_omega_nu = no
fit_ns = yes
fit_w0 = yes
fit_wa = yes
.....
#-----
perturb_type = noperturbsup
gauge = quintsynchronous
.....
# Equation of state
w0      = -1.095936698
min_w0  = -3.
max_w0  = -0.3
error_w0 = 0.2
# Equation of state variation.
wa      = -0.2295558424
min_wa  = -2.5
max_wa  = 2.5
error_wa = 0.8

```

- **Format totalement libre (fichier texte)**
- **Héritage des paramètres globaux**

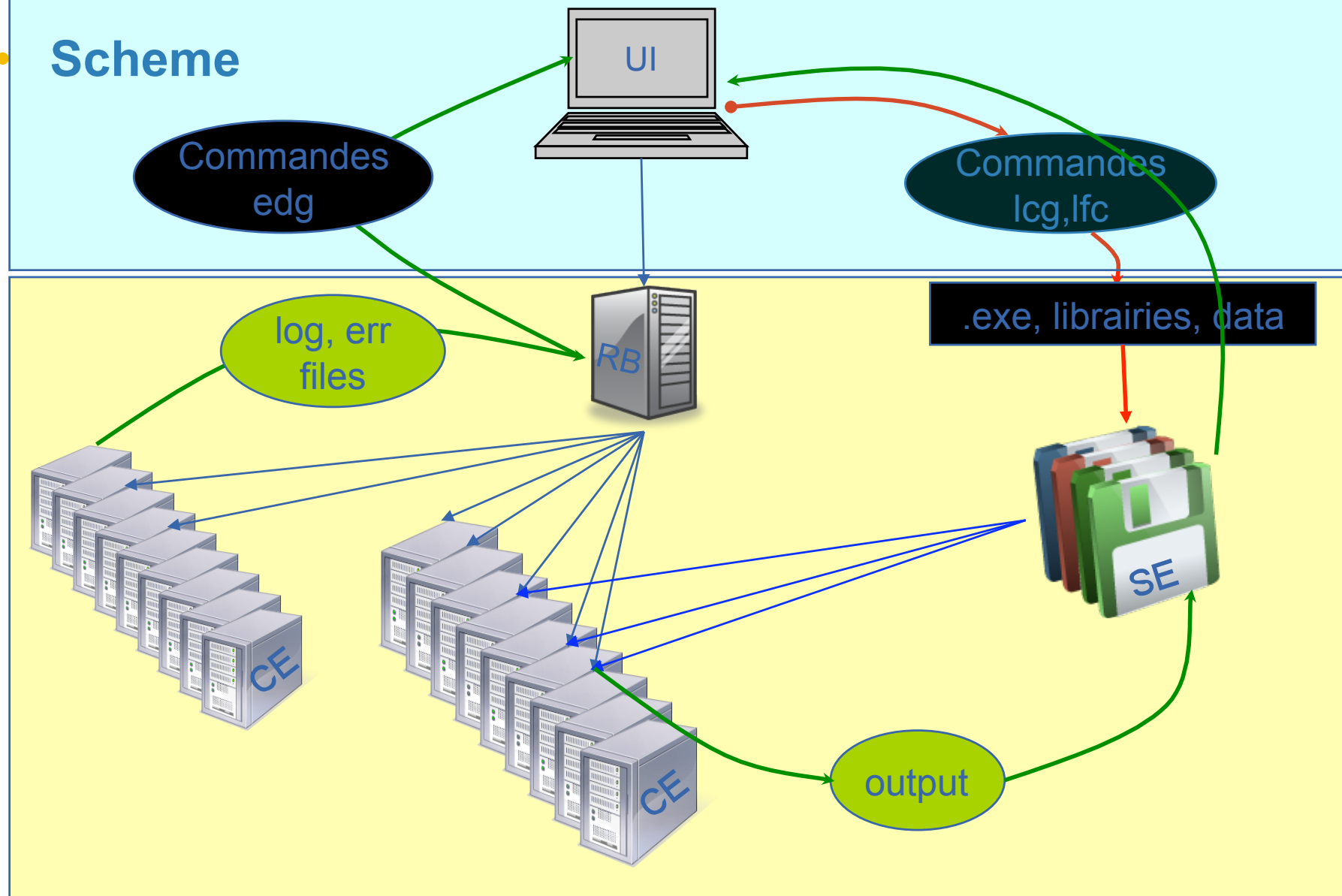
```
#      General configuration file

.....
#-----
perturb_type = noperturbsup
gauge = quintsynchronous
.....
# Equation of state
w0      = -1.095936698
min_w0  = -3.
max_w0  = -0.3
error_w0 = 0.2
# Equation of state variation.
wa      = -0.2295558424
.....
```

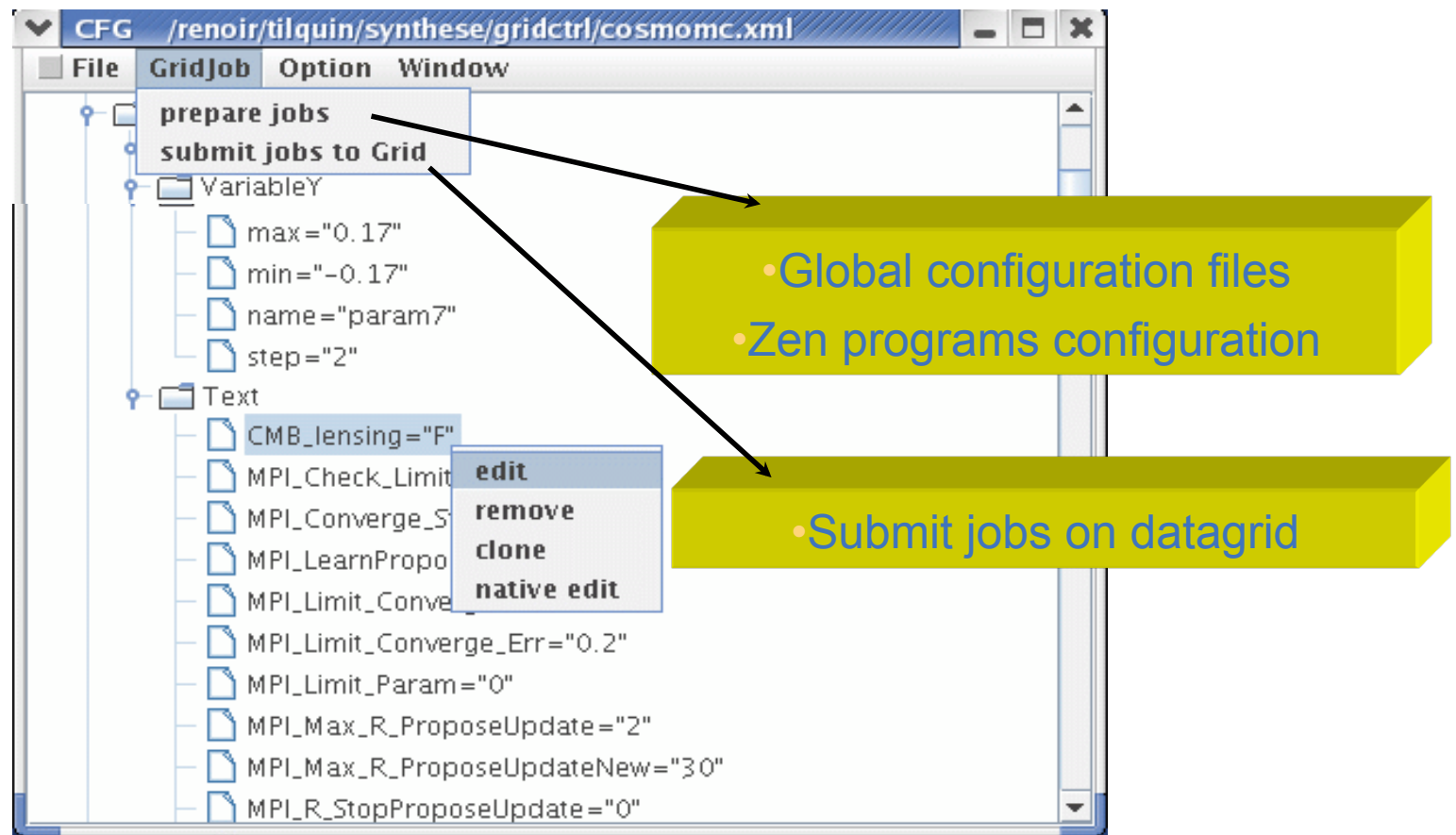
configurateur

```
#      Configuration file for BAO
omegab = [omega_b]
omegam = [omegam]
h = [h]
ns = [ns]
tau = [tau]
w0 => [w0]
wa => [wa]
sigma8 = [sigma8]
bias = [bao_bias]
#-----
Internal_flag = 1
#      Should be zero. Curvature not
#      yet implemented in bao
omegak = 0
.....
```

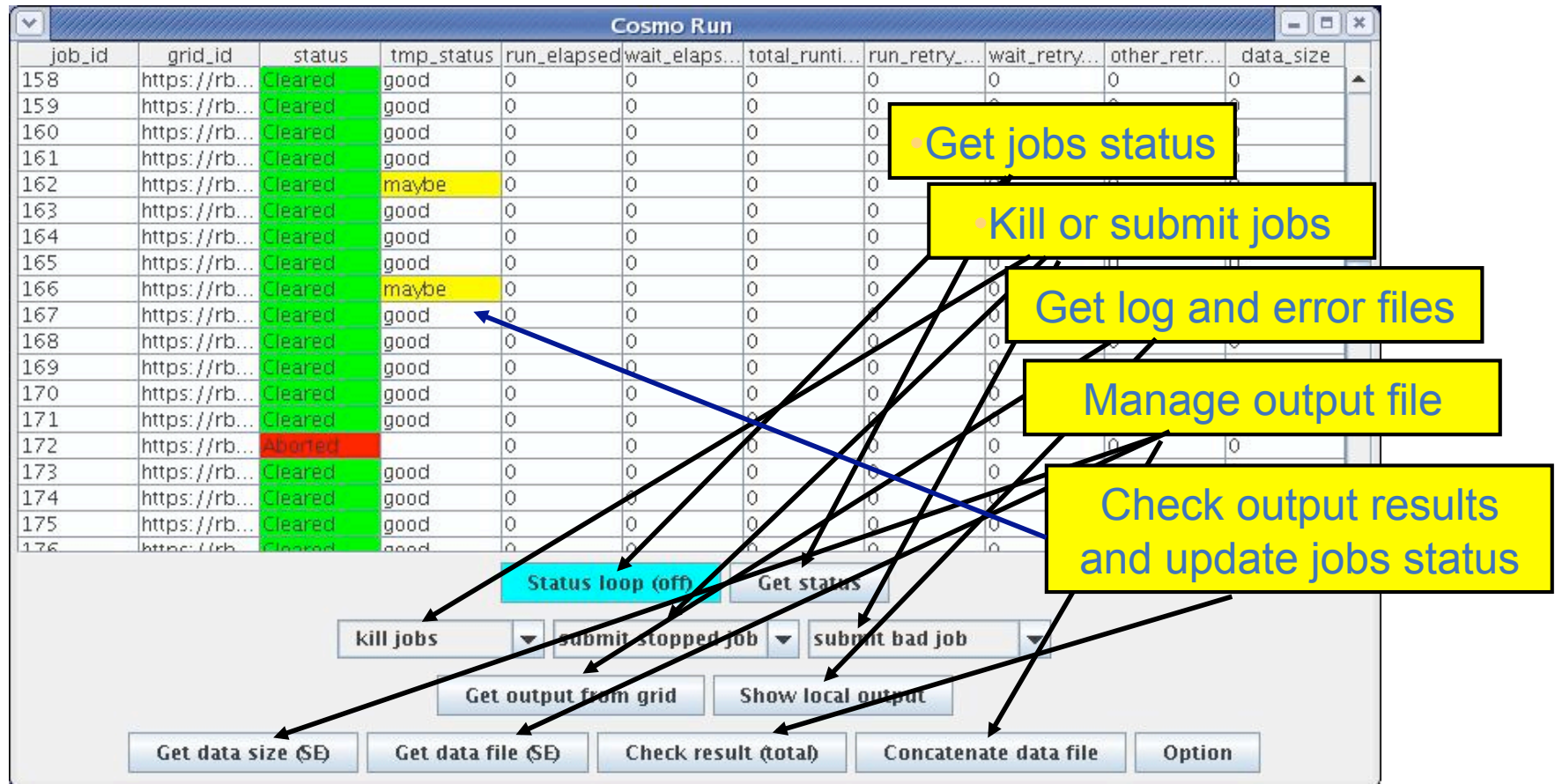
Scheme



- A run is a set of $n \times n$ jobs (400) send in parallel. A graphical interface has been developed (thanks to Zuxuan Qian) to construct and submit them.



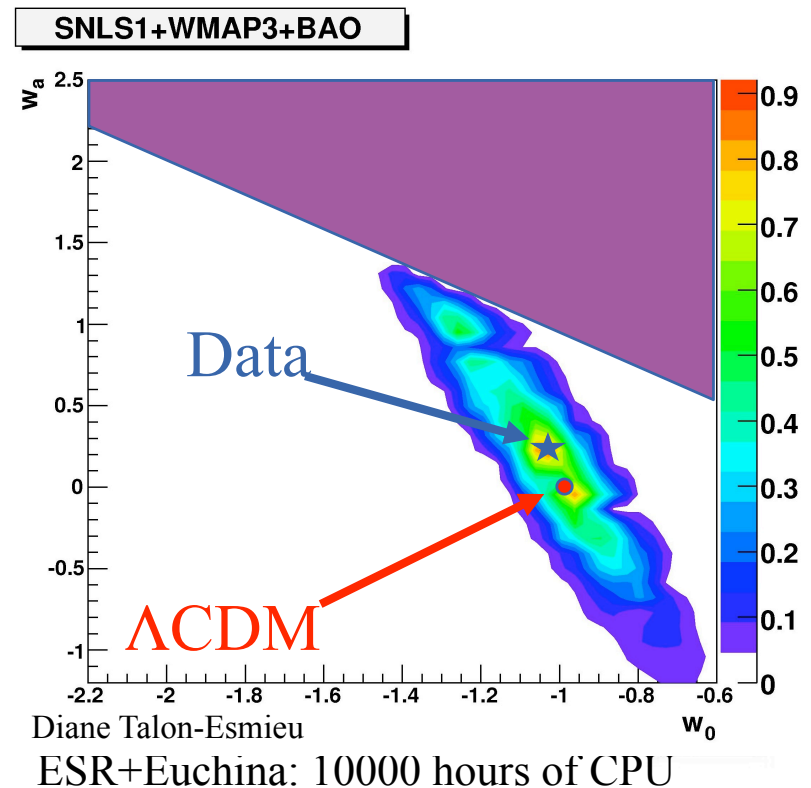
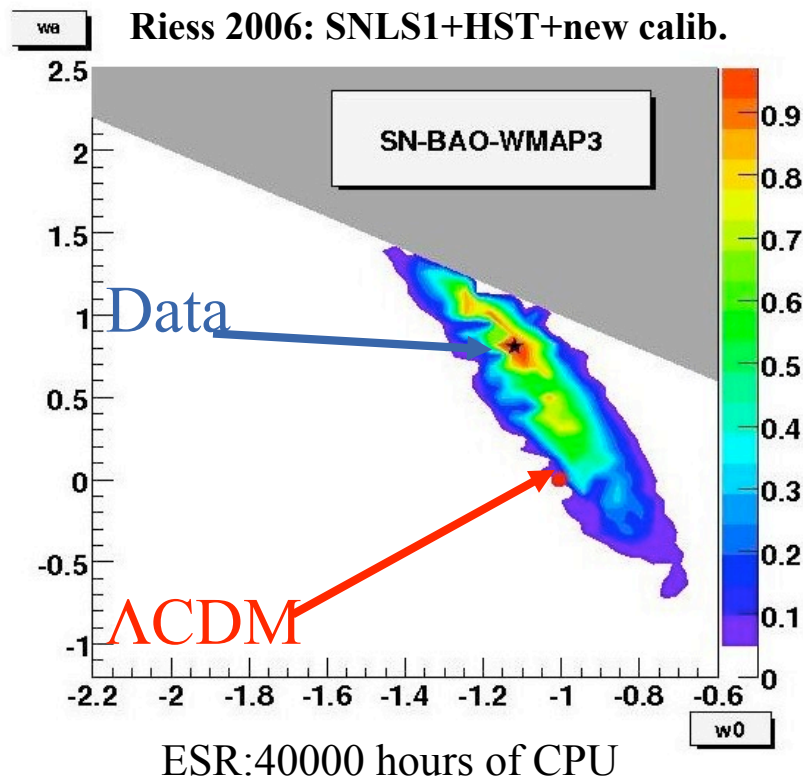
- Jobs monitoring and output data manager



The screenshot shows the 'Cosmo Run' graphical interface. It features a table with columns: job_id, grid_id, status, tmp_status, run_elapsed, wait_elaps..., total_runti..., run_retry..., wait_retry..., other_retr..., and data_size. The table lists jobs 158 through 176. Job 172 is highlighted in red with the status 'Aborted', while others are green ('Cleared') or yellow ('maybe'). Below the table are several control buttons: 'Status loop (off)', 'Get status', 'kill jobs', 'submit stopped job', 'submit bad job', 'Get output from grid', 'Show local output', 'Get data size (\$E)', 'Get data file (\$E)', 'Check result (total)', 'Concatenate data file', and 'Option'. Five yellow callout boxes with arrows point to specific parts of the interface:

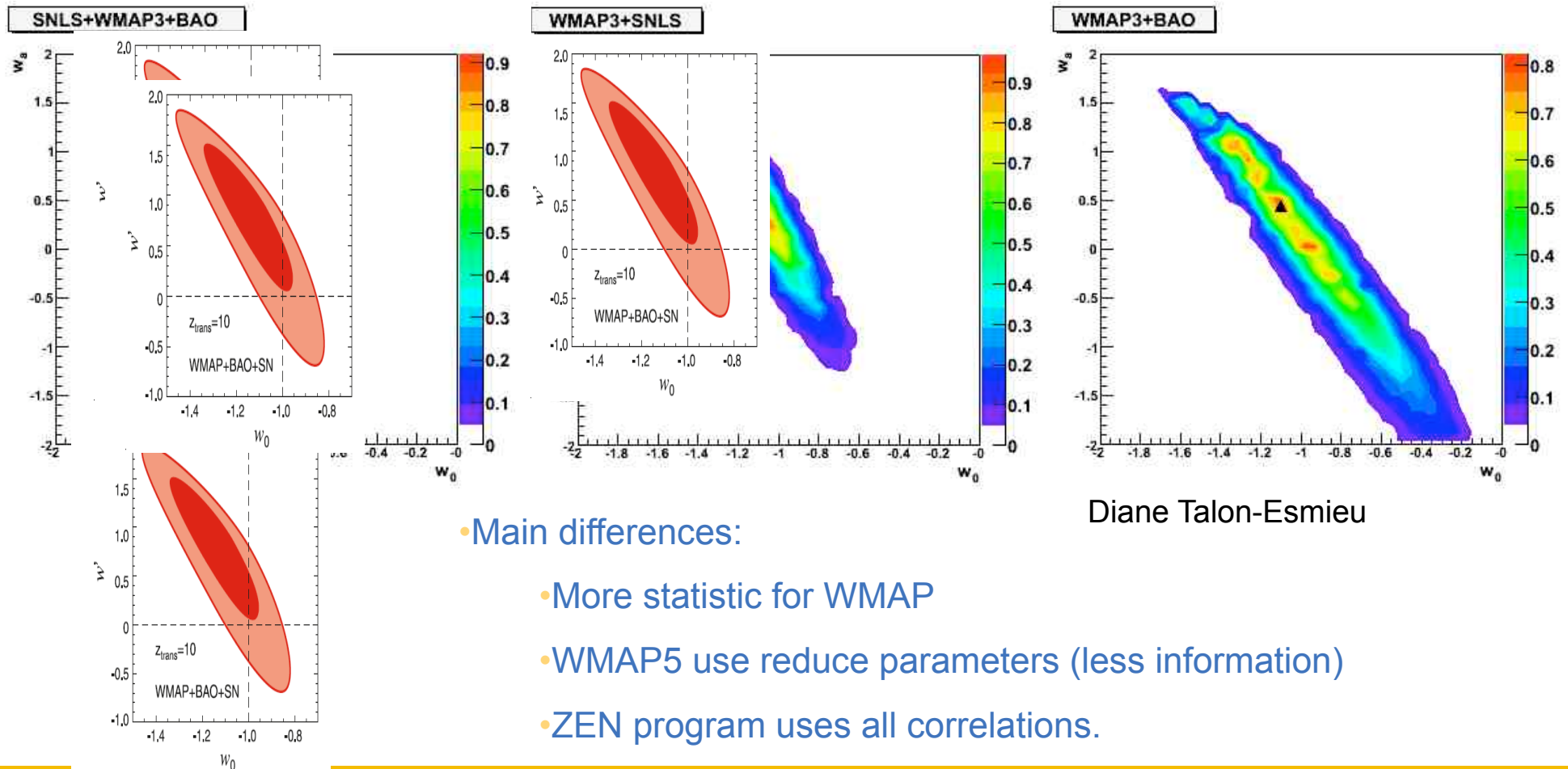
- 'Get jobs status' points to the 'Get status' button.
- 'Kill or submit jobs' points to the 'kill jobs', 'submit stopped job', and 'submit bad job' buttons.
- 'Get log and error files' points to the 'Get output from grid' button.
- 'Manage output file' points to the 'Show local output' button.
- 'Check output results and update jobs status' points to the 'Check result (total)' button.

- **First results using datagrid within:**
 - ESR (Earth Science Research) VO (thanks to M. Petit-Didier) and Euchina Virtual organization
 - Using SN+CMB+BAO with frequentist statistic and (BAOfit from Sun Lei, PKU-CPPM Join PhD)



- News results on WMAP3-SNLS-BAO comparison

Comparison with WMAP5+SNLS+BAO last publication



Diane Talon-Esmieu

- Main differences:
 - More statistic for WMAP
 - WMAP5 use reduce parameters (less information)
 - ZEN program uses all correlations.

- **The ZEN program is now running on datagrid**
Thanks to ESR virtual organization
- **The two graphical interfaces are very powerful tools to submit jobs and to debug.**
- **Major problem was the linux version (SL3,SL4)**
- **Our first results on multi-probes analysis are compatible with competitive analysis (Bayesian) and WMAP5**
- **Perspectives:**
 - Add new probes (WL, etc...) and add new data (WMAP5,SNLS3...)
 - Investigate new theoretical models and implement them in ZEN
 - Add new statistics (MCMC, Bayesian...)
 - Start systematic analysis !
 - Write a documentation : IT IS A PUBLIC TOOL
 - If you want to use it you are welcome (possible training).